

# The eruption of 22 April 2021 as observed by Solar Orbiter, STEREO and Earth bound instruments

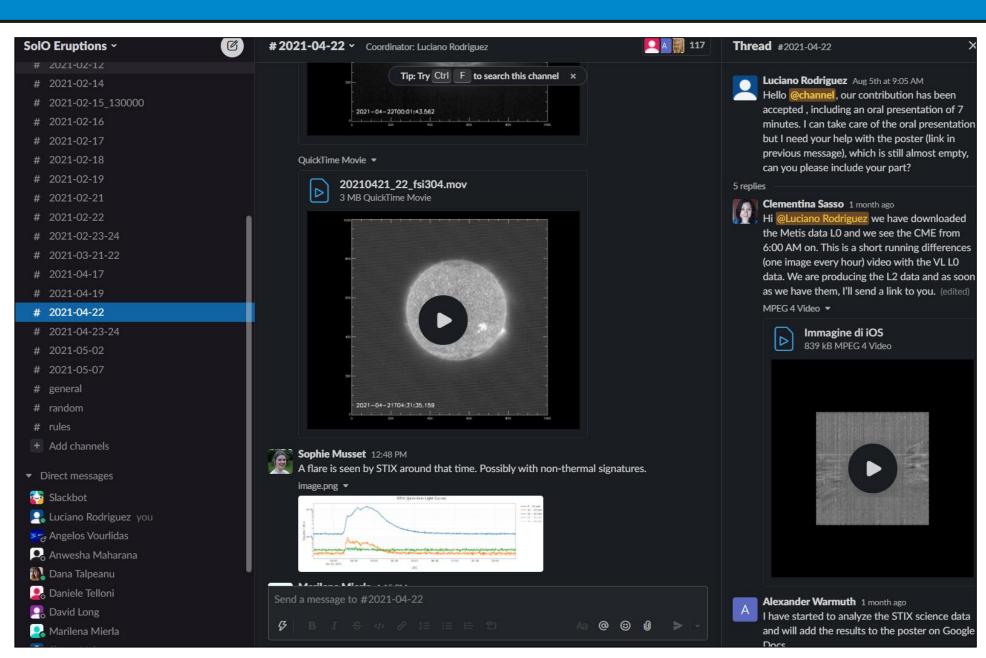
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#### Ongoing work, started on Slack

- Solar Orbiter Eruptive Events Remote Sensing Science Working Group
- Thanks to the admins: Stephanie Yardley and Kévin Dalmasse.





#### **Abstract**

The Extreme Ultraviolet Imager (EUI) onboard Solar Orbiter (SolO) observed an eruption with its Full Sun Imager (FSI) in both of its channels (17.4/30.4 nm), on 2021-April-22. At the time, the spacecraft was at 0.87 au from the Sun. The eruption was seen at the SW limb, starting at 04:24 UT, with the source slightly backside as seen from SolO (S20W103).

From the Earth's perspective, SDO/AIA and PROBA2/SWAP observed a wave and dimmings starting around 04:07 UT, on-disk at ~S20W05.

STEREO-A/EUVI saw similar signatures of an eruption starting around 04:17 UT, on-disk at ~S20W50. T

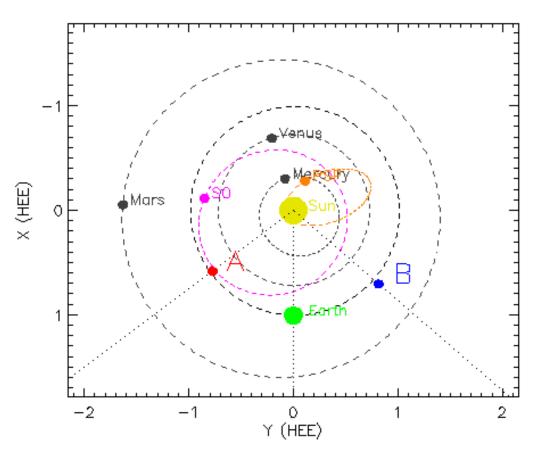
he corresponding CME was visible shortly after in several coronagraphs. SOHO/LASCO-C2 observed a full halo CME starting around 06:00 UT. STEREO-A/COR2 recorded a clear structured CME seen from around 05:23 UT. SolO Metis data will be analysed as it becomes available.

SolO/STIX observed the associated long-duration X-ray flare, which was partially occulted. This allows the characterization of both the thermal plasma and any potential contribution of nonthermal electrons in the tenuous coronal source. The full-Sun radiometer PROBA2/LYRA also observed the event with the two SXR channels.

The corresponding ICME arrived at the Earth on 2021-April-24-25 (probably also at STEREO-A), it was driving a shock and created minor geomagnetic storm conditions. STEREO-A/SEPT and ACE/EPAM observed a weak particle event most likely related to this eruption. We will analyse in depth these CME-ICME connections.



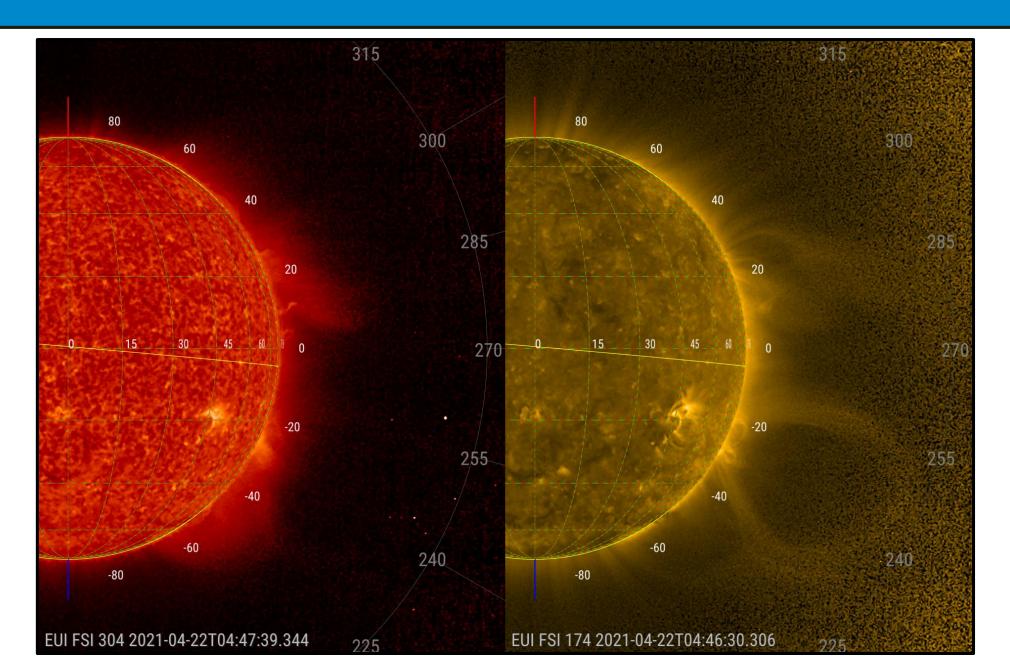
#### Spacecraft positions



	Date [yyyy-mm- dd]	Sep. angle SOLO - Earth [deg]	Sep. angle SOLO - STA [deg]	lon_SOLO HEEQ [deg]	lat_SOLO HEEQ [deg]	dist_SOLO _Sun [AU]
	2021-04-22	97.96	45.05	-97.97	0.32	0.86



#### The CME in SolO EUI observations





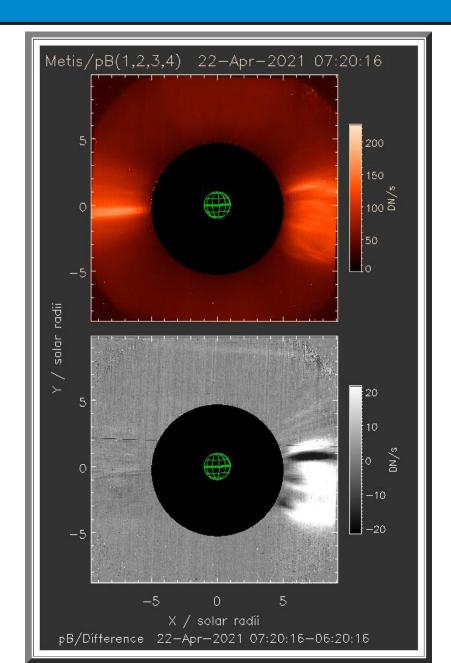
#### SolO EUI 174



- FSI synoptic program
- ~5 minute cadence



#### SolO Metis

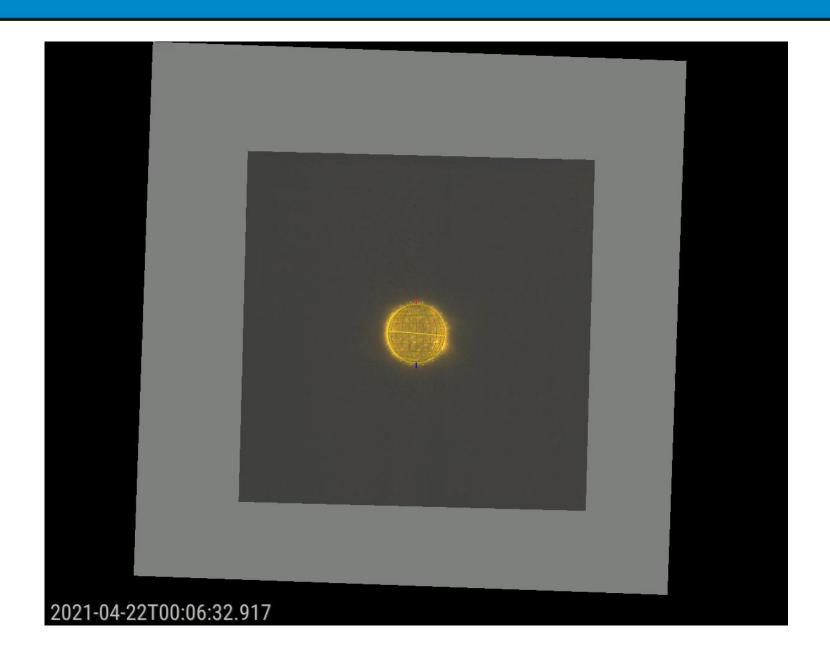


- On April 22, Metis was running a low-frequency synoptic program.
- Cadence of 1 hour
- 4 VL images to form a single pB image

#### Movie

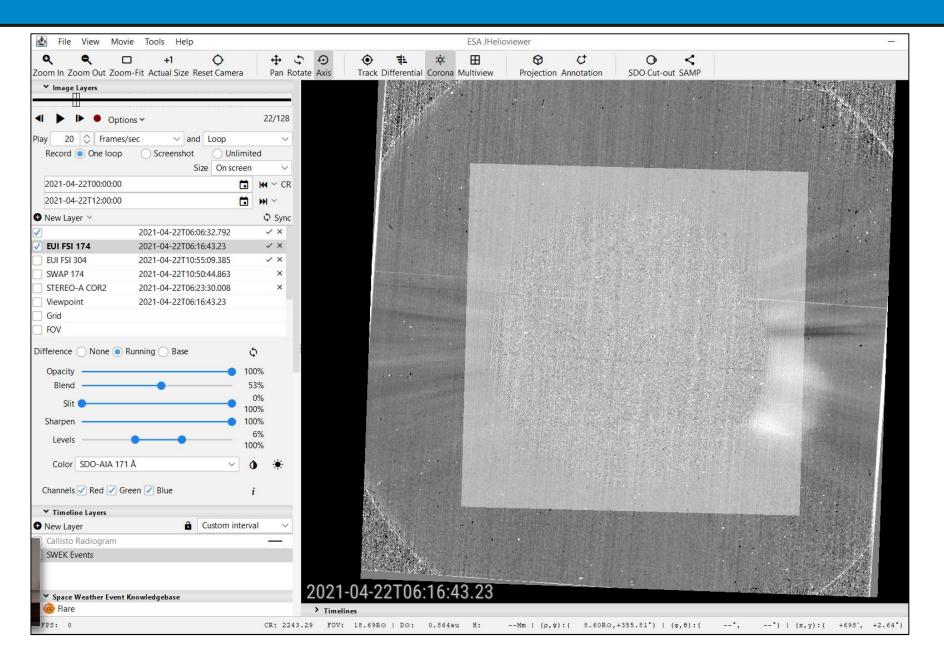


#### **SolO EUI-Metis**





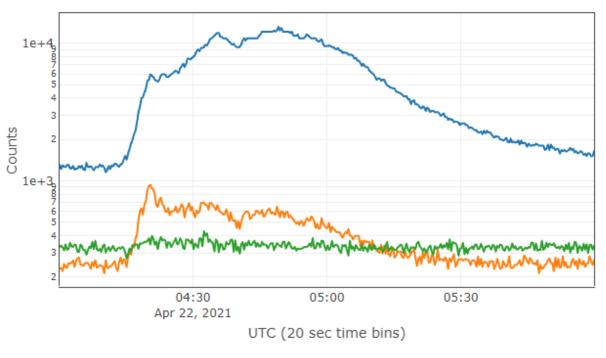
#### SolO in JHV

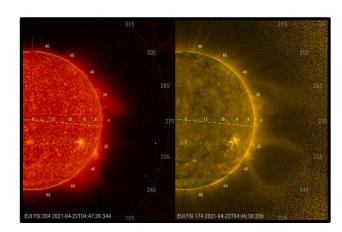




#### SolO STIX







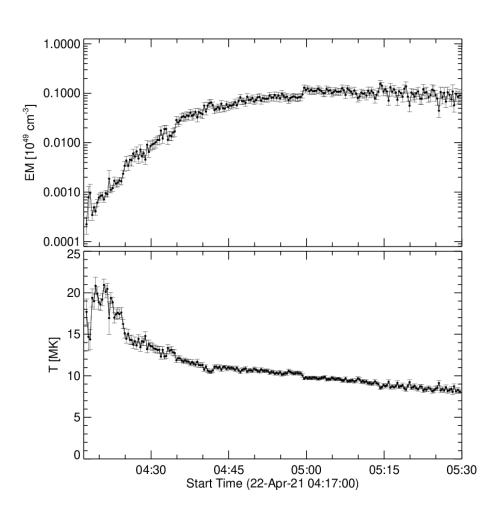
4 – 10 keV

10 - 15 keV 15 - 25 keV 25 - 50 keV 50 - 84 keV

- STIX X-ray light curves of the associated flare (GOES class C3.8)
- Long-duration event with an extended decay phase.
- No pronounced emission above 20 keV, indicative of a weak nonthermal component.
- The flare is partially occulted: chromospheric footpoints which usually dominate the emission are behind the solar limb.



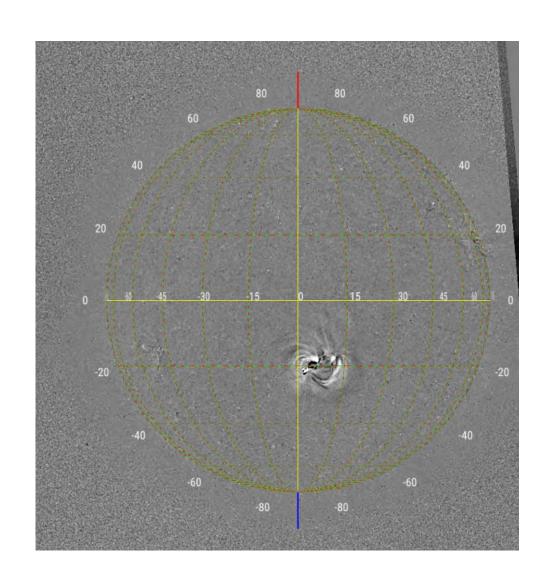
#### SolO STIX

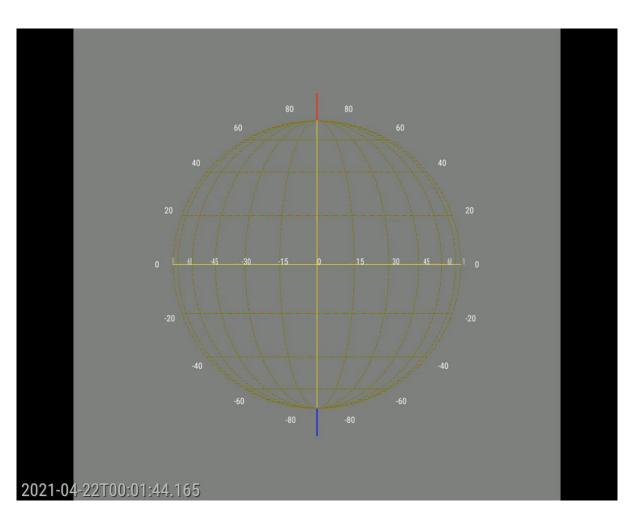


- Evolution of emission measure and temperature of the hot plasma (from X-ray spectra).
- The impulsive phase shows rapid heating to 20 MK, followed by gradual cooling. Peak thermal energy of  $2.4\pm0.3x10^{29}$  erg.



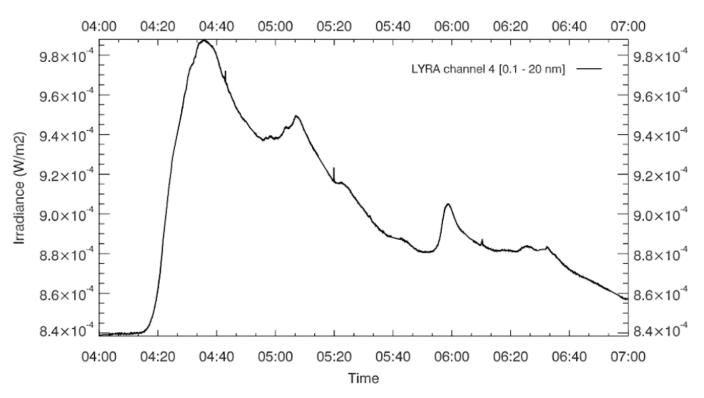
#### PROBA2- SWAP







#### PROBA2-Lyra

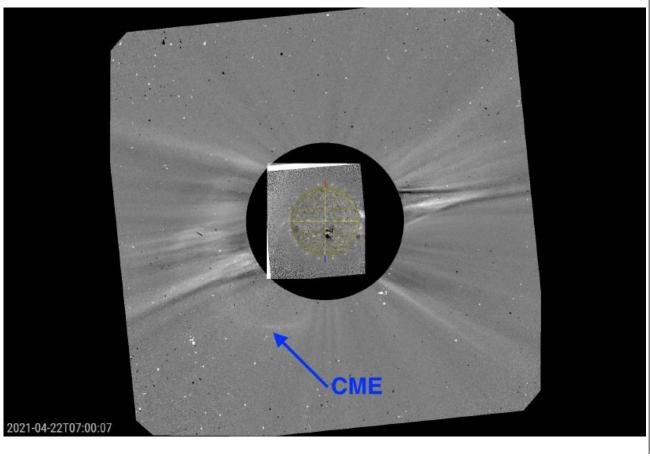


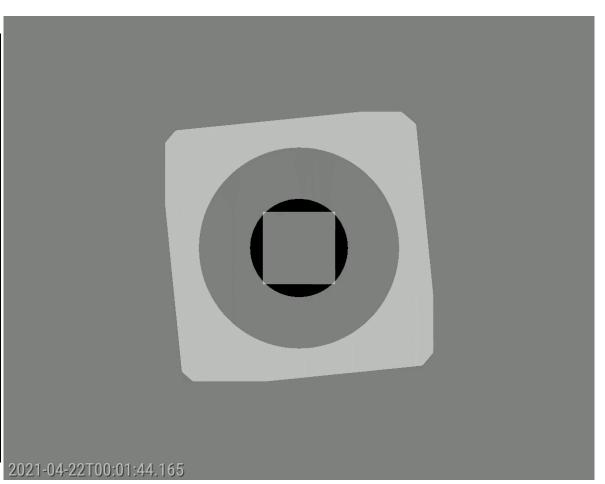
- The C3.8 flare associated to the CME, as seen by PROBA2/LYRA channel 4 (0.1-20 nm).
- The active region was close to the Sun-center. No occultation here, contrarily to Solar Orbiter.
- Long duration event, associated multiple episodes of plasma heating, following a brief main impulsive phase.
- Not all of these secondary heating episodes were produced by the same active regions. Another active region located close to the west limb was, produced the two peaks after 06:00 AM.



#### SOHO-LASCO C2

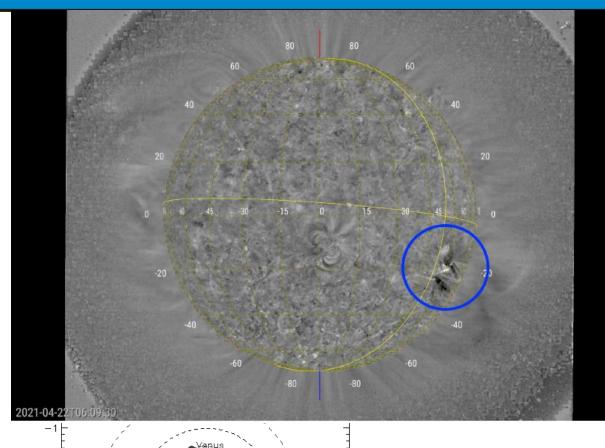
- Weak partial halo
- $V = ^350 \text{ km/s}$

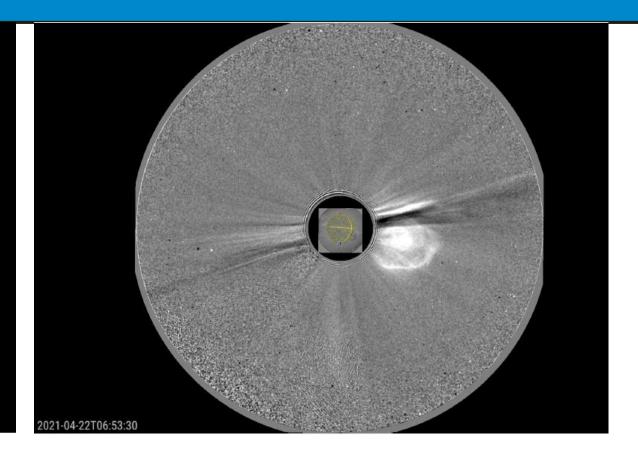


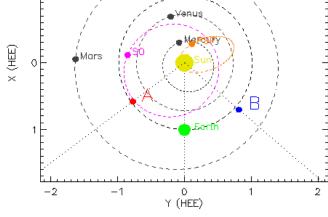




#### STEREO-A EUVI and COR2

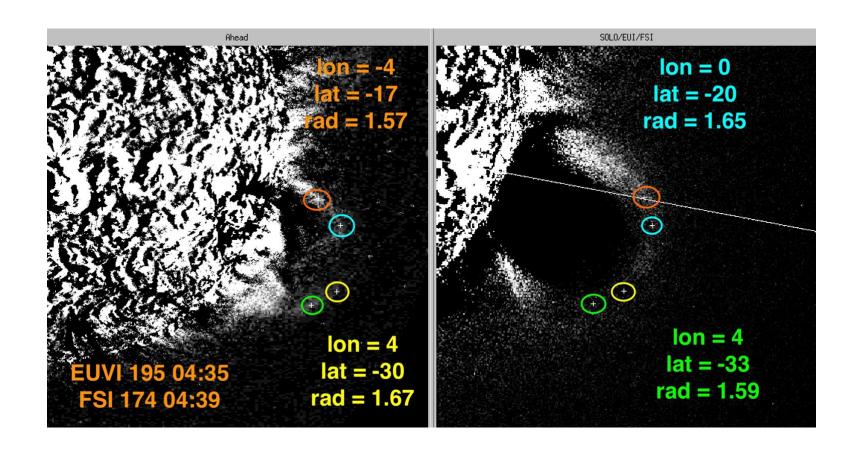








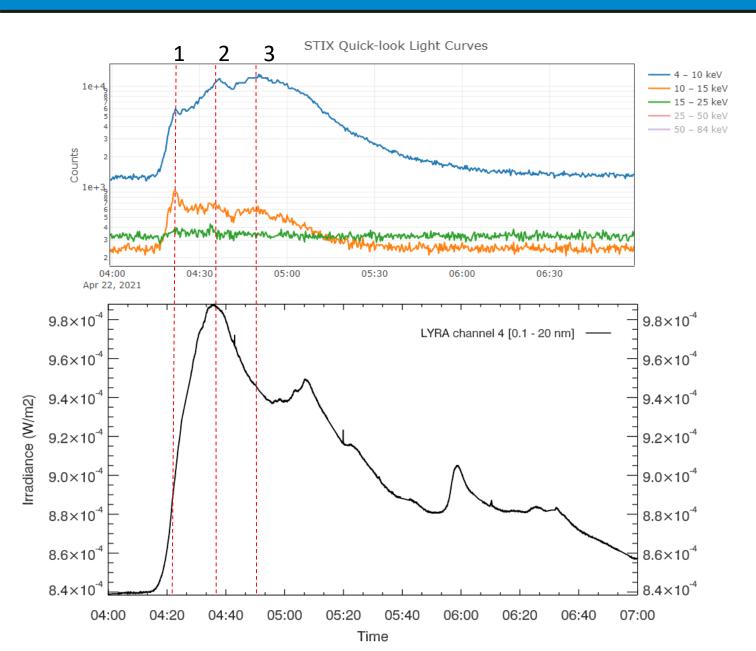
#### Triangulation, FSI174 + EUVI-A 195



- Longitudinal values of the triangulation are consistent with an event occurring close to disk center as seen from the Earth, and at the limb as seen from SolO
- Coronagraph triangulation was not accurate



#### Comparison SolO STIX - PROBA2 Lyra



Each peak in the STIX thermal X-ray emission reflects a heating event.

Correlation with UV emission:

- 1. in steeply rising phase of UV
- 2. around peak of UV
- 3. after peak of UV (small bump)

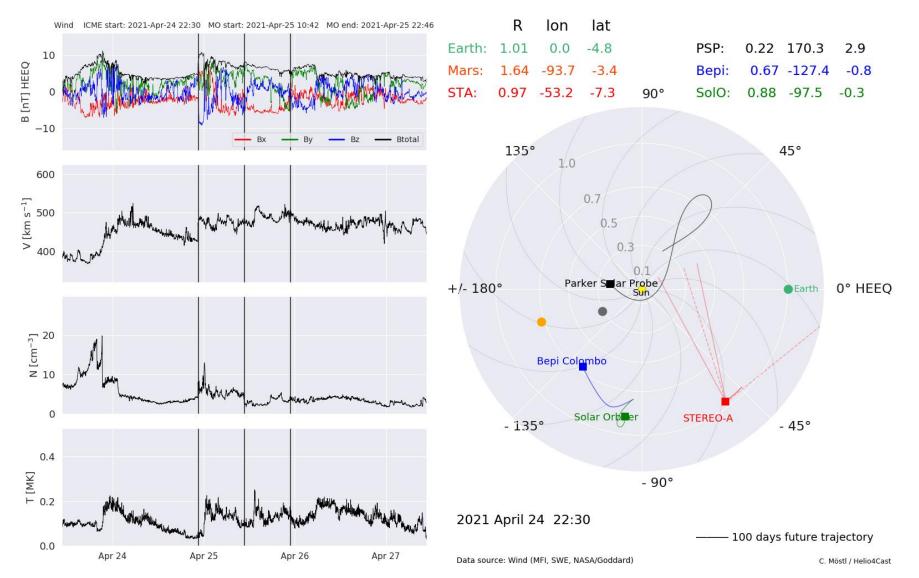
Later UV peaks are not seen in Xrays –from different locations fully occulted for STIX

Main peak (2):

X-rays are delayed by ~100 s with respect to UV – typical timesale for chromospheric evaporation

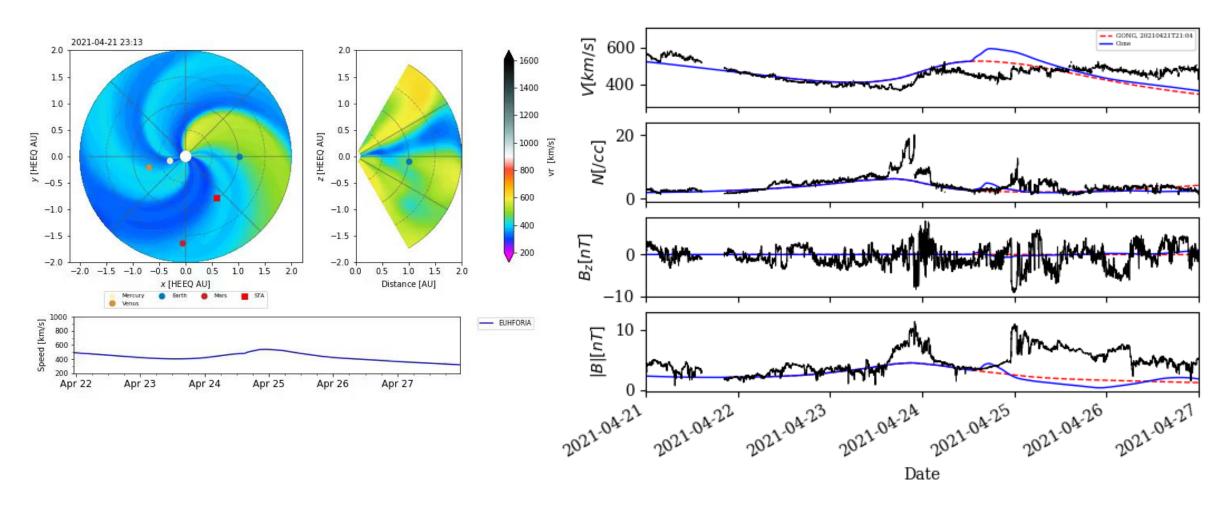


## ICME in situ





### ICME in situ - EUHFORIA





#### **Summary/Conclusions**

- CME from 22 April 2021, seen by SolO, SOHO, STEREO-A, PROBA2
- The flare was observed as a partially occulted event by STIX and fully visible by Lyra
- Triangulation between EUVI and EUI could be done, but not for the coronagraphs
- The ICME arrived to the Earth late on 24 April
- The ICME can be simulated with EUHFORIA relatively well





# EXTRA



#### Solo EUI 304



